

# Experimental estimation of one-parameter qubit gates in the presence of phase diffusion

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We address estimation of one-parameter qubit gates in the presence of phase diffusion. We evaluate the ultimate quantum limits to precision, seek for optimal probes and measurements, and demonstrate an optimal estimation scheme for polarization encoded optical qubits. An adaptive method to achieve optimal estimation in any working regime is also analyzed in details and experimentally implemented.

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## I. INTRODUCTION

Suppose you are given a black box which operates on qubits. The reconstruction of the corresponding quantum operations [1, 2, 3] is critical to verify its actions as a quantum logic gate [4] as well as to characterize decoherence processes [5]. Let us consider the case when the action of the device is described by the unitary  $U_\phi = \exp\{-i\sigma_n\phi\}$  and corresponds to a phase-shift (rotation)  $\phi$  about a known axis  $n$ . This is the simplest operation on a qubit and realizes a one-parameter logical gate, which allows, combined with the Hadamard gate, the transformation of any qubit state into another. The quantum characterization [6, 7, 8, 9, 10, 11, 12] of this kind of devices is of interest for quantum information processing and amounts to estimate the phase-shift by measuring a suitable observable at the output. In the following, we fix the reference frame and assume, without loss of generality, rotations  $U_\phi = \exp\{-i\sigma_z\phi\}$  about the  $z$  axis.

In ideal conditions the estimation of the phase-shift consists of preparing a qubit in a known pure state  $\varrho = |\psi_0\rangle\langle\psi_0|$  and then performing a suitable measurements on the (pure) shifted state  $\varrho_\phi = U_\phi\varrho U_\phi^\dagger$ . In a realistic implementation, however, the propagation of a qubit is unavoidably accompanied by some noise, which influences the estimation scheme and usually degrades the overall precision. In this paper we address estimation in the presence of the most detrimental kind of noise for a phase gate, i.e. non dissipative phase noise, which destroys the off diagonal elements of the density matrix and thus the information on the imposed phase-shift. We evaluate the ultimate quantum limits to precision in the presence of noise and determine both the optimal preparation of the probe qubit and the optimal measurement to be performed at the output. The optimal estimation scheme is then experimentally demonstrated for polarization encoded optical qubits, together with an adaptive method to achieve optimal estimation in any working regime.

The paper is structured as follows. In section II we describe the system under investigation focusing on the ultimate limits to precision of phase-shift estimation in the presence of phase noise. The quantum Cramér-Rao limit as well as the optimal quantum estimator are explicitly given. Section III introduces a realistic scenario for qubit phase-shift estimation: an estimation scheme based on spin measurements is described in detail and the corresponding Fisher information is derived. In section IV we turn our attention to optical qubit systems and describe phase-shift estimation under two different approaches: the inversion method and the Bayesian estimation. We also investigate numerically the robustness of the Bayesian analysis w.r.t. the inversion method in the non asymptotic regime, i.e., in the case of small amounts of data. Section V is devoted to the experimental demonstration of the optimal scheme for phase-shift estimation, and of the adaptive method to achieve the quantum Cramér-Rao bound in any working regime. Section VI closes the paper with some concluding remarks.

## II. PHASE-SHIFT ESTIMATION IN QUBIT SYSTEMS

The measurement scheme we are going to address is schematically depicted in Fig. 1.

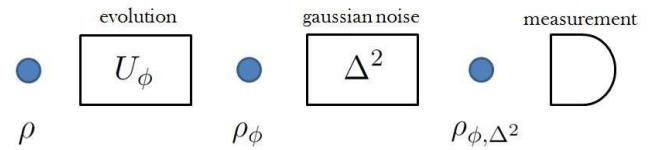


FIG. 1: Schematic diagram of the measurement scheme.

One has a single qubit, initially prepared in the pure state  $\varrho = |\psi_0\rangle\langle\psi_0|$ , that undergoes an unknown phase-shift  $\phi$  imposed by the unitary  $U_\phi$ . Before being measured the shifted

state  $\varrho_\phi = U_\phi \varrho U_\phi^\dagger$  is degraded by a non dissipative phase noise occurring during the propagation. The effect of this kind of noise on the qubit density matrix, given a noise factor  $\gamma$ , can be described by the following Master equation (ME):

$$\dot{\varrho}_{\phi, \Delta^2} = \gamma \mathcal{L}[\sigma_+ \sigma_-] \varrho_{\phi, \Delta^2}, \quad (1)$$

where  $\mathcal{L}[A] \varrho_{\phi, \Delta^2} = \frac{1}{2} \{ [A \varrho_{\phi, \Delta^2}, A^\dagger] + [A, \varrho_{\phi, \Delta^2} A^\dagger] \}$ , and  $\Delta^2 = \gamma t/2$  is, as we will see in the following lines, the effective noise factor. Since  $\mathcal{L}[\sigma_+ \sigma_-]$  and  $\sigma_z$  commute, we can focus on the evolution of  $\varrho$ , i.e.,  $\dot{\varrho}_{\Delta^2} = \gamma \mathcal{L}[\sigma_+ \sigma_-] \varrho_{\Delta^2}$ . Upon writing  $\varrho_{\Delta^2}$  in the eigenbasis of  $\sigma_z$ , the ME leads to differential equations for the matrix elements  $\varrho_{nm}(t) = \langle n | \varrho_{\Delta^2} | m \rangle$ , where  $\dot{\varrho}_{nm}(t) = -\frac{1}{2} \gamma (n - m)^2 \varrho_{nm}(t)$  whose solutions read:

$$\varrho_{nm}(t) = e^{-\Delta^2(n-m)^2} \varrho_{nm}(0). \quad (2)$$

where  $\varrho_{nm}(0)$  are the initial density matrix elements. From Eq. (2) it is clear that, whereas the diagonal elements are left unchanged by the evolution under ME (1), and, in turn, energy is conserved, the off-diagonal ones are progressively destroyed. Finally, the solution of Eq. (1) is  $\varrho_{\phi, \Delta^2} = U_\phi \varrho_{\Delta^2} U_\phi^\dagger$ . Since we can consider the noise factor  $\Delta^2$  as a fixed parameter, in the following we will not write explicitly the dependence on it.

It is worth noting that the same evolution as (2) can be also obtained by the application of a random, zero-mean Gaussian-distributed phase-shift to a quantum state. Since the phase shift of an amount  $\varphi$  is described by the unitary operator  $U_\varphi \equiv \exp(-i\varphi \sigma_z)$ , we can write the state degraded by the Gaussian phase noise as follows:

$$\varrho_{Gn} = \int_{\mathbb{R}} d\varphi \frac{e^{-\varphi^2/(4\Delta^2)}}{\sqrt{4\pi\Delta^2}} U_\varphi \varrho(0) U_\varphi^\dagger \quad (3)$$

$$= \sum_{nm} \int_{\mathbb{R}} d\varphi \frac{e^{-\varphi^2/(4\Delta^2)}}{\sqrt{4\pi\Delta^2}} e^{-i\varphi(n-m)} \varrho_{nm}(0) |n\rangle \langle m| \quad (4)$$

$$= \sum_{nm} e^{-\Delta^2(n-m)^2} \varrho_{nm}(0) |n\rangle \langle m|, \quad (5)$$

which is the same as in Eq. (2). This point will be useful for the experimental demonstration.

The goal of an estimation problem is not only retrieve the actual value of the unknown parameter (the phase shift in our case), but obtain this information with the minimum uncertainty. The ultimate limit to the precision one can reach, i.e., the minimum variance, is given by the quantum Cramér-Rao bound [13, 14, 15, 16]:

$$\text{Var}[\phi] = [N H(\phi)]^{-1}, \quad (6)$$

where  $N$  is the number of measurements and  $H(\phi)$  is the quantum Fisher information (QFI). If we choose a pure probe state,  $\varrho = |\psi_0\rangle \langle \psi_0|$ , then the QFI equals four times the fluctuation of  $\sigma_z$  i.e one simply has  $H = 4(1 - \langle \psi_0 | \sigma_z | \psi_0 \rangle^2)$ . More in general, for mixed states,  $H(\phi)$  can be written as [17]:

$$H(\phi) = 2 \sum_{n \neq m} \frac{(\lambda_n - \lambda_m)^2}{\lambda_n + \lambda_m} |\langle \psi_m(\phi) | \partial_\phi \psi_n(\phi) \rangle|^2, \quad (7)$$

where the  $|\psi_n(\phi)\rangle$ 's are the eigenvectors of the state  $\varrho_\phi = U_\phi \varrho U_\phi^\dagger$  and  $\lambda_n$  the corresponding eigenvalues.

The spectral decomposition of  $\varrho_\phi$  reads as follows

$$\varrho_\phi = \sum_n \lambda_n |\psi_n(\phi)\rangle \langle \psi_n(\phi)| = \sum_n \lambda_n U_\phi |\psi_n\rangle \langle \psi_n| U_\phi^\dagger. \quad (8)$$

$|\psi_n\rangle$  being the eigenvectors of the initial state. If we decompose  $|\psi_n(\phi)\rangle$  in the standard basis as follows:

$$|\psi_n(\phi)\rangle = U_\phi |\psi_n\rangle = U_\phi \sum_k r_{nk} |k\rangle, \quad (9)$$

then, by substituting (9) into the eigenvalues equation  $\varrho_\phi |\psi_n(\phi)\rangle = \lambda_n |\psi_n(\phi)\rangle$ , after some algebra we obtain:

$$\sum_k \varrho_{nk}(0) e^{-\Delta^2(n-k)^2} r_{qk} = \lambda_q r_{qn} \quad \forall n. \quad (10)$$

Moreover, since  $|\partial_\phi \psi_n(\phi)\rangle = i \sum_k k r_{nk} e^{ik\phi} |k\rangle$ , we have [see Eq. (7)]:

$$|\langle \psi_m(\phi) | \partial_\phi \psi_n(\phi) \rangle|^2 = \left| \sum_k k r_{mk} r_{nk} \right|^2. \quad (11)$$

Finally, given  $\lambda_n$  and  $r_{nk}$ , we can evaluate the QFI, which is, as expected for a shift parameter, independent of  $\phi$ .

Generally, the calculation of the eigenvalues  $\lambda_n$  and the coefficients  $r_{nk}$  of the standard basis decomposition (9) is a difficult task, which can be performed by means of numerical analysis. However, in the case of qubit systems, all the calculations can be carried out analytically, as we are going to show in the following.

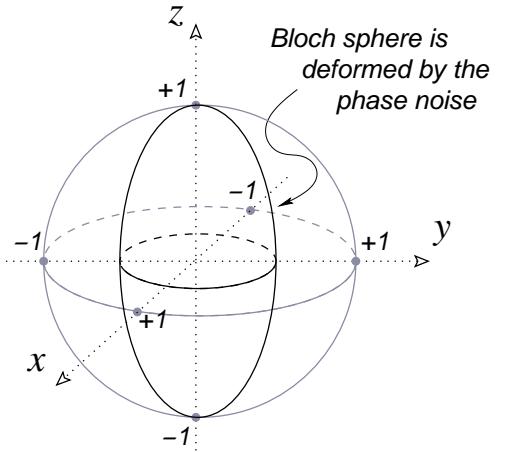


FIG. 2: Effect of the phase destroying ME on the Bloch sphere: the  $z$  component is left unchanged while  $x$  and  $y$  ones are scaled by a factor  $e^{-\Delta^2}$ . See the text for details.

Upon writing the initial qubit state in the Bloch sphere representation

$$\varrho = \frac{\mathbb{1} + \mathbf{r} \cdot \boldsymbol{\sigma}}{2}, \quad (12)$$

where  $\mathbb{I}$  is the  $2 \times 2$  identity matrix,  $\mathbf{r} = (r_x, r_y, r_z)$ ,  $|\mathbf{r}|^2 \leq 1$ , and  $\boldsymbol{\sigma} = (\sigma_x, \sigma_y, \sigma_z)$  is the Pauli matrices vector, then its evolution under the action of the ME (1) can be reduced to the following transformation of the Bloch vector  $\mathbf{r}$  [18]:

$$(r_x, r_y, r_z) \rightarrow (r_x e^{-\Delta^2}, r_y e^{-\Delta^2}, r_z), \quad (13)$$

i.e., the Bloch sphere is deformed in such a way that the  $z$  component is left unchanged while  $x$  and  $y$  ones are scaled by a factor  $e^{-\Delta^2}$  (see Fig. 2). Now, due to symmetry considerations, without lack of generality we will focus our attention on the pure state with Bloch vector:

$$\mathbf{r} = (\sin 2\theta, 0, \cos 2\theta), \quad (14)$$

$2\theta$  being the azimuthal angle ( $\theta = 0$  and  $\theta = \pi/2$  correspond to the north and south poles of the Bloch sphere, respectively).

In the density matrix representation (choosing the  $\sigma_z$  eigenvectors basis), after the phase-noise evolution, we have:

$$\varrho = \begin{pmatrix} \cos^2 \theta & e^{-\Delta^2} \cos \theta \sin \theta \\ e^{-\Delta^2} \cos \theta \sin \theta & \sin^2 \theta \end{pmatrix}. \quad (15)$$

The two eigenvalues are:

$$\lambda_{\pm} = \frac{1}{2} \left( 1 \pm \frac{1 + f(\theta, \Delta^2)}{\sqrt{2}} \right), \quad (16)$$

with  $f(\theta, \Delta^2) = \sqrt{e^{-2\Delta^2} + (1 - e^{-2\Delta^2}) \cos 4\theta}$ , and the corresponding eigenvectors read:

$$|\psi_{\pm}\rangle = \frac{1}{Z_{\pm}} \begin{pmatrix} g_{\pm}(\theta, \Delta^2) \\ 1 \end{pmatrix}, \quad (17)$$

$$= \frac{1}{Z_{\pm}} [g_{\pm}(\theta, \Delta^2)| + 1/2\rangle + | - 1/2\rangle], \quad (18)$$

$\langle \psi_+ | \psi_- \rangle = 0$ , and:

$$g_{\pm}(\theta, \Delta^2) = \cos 2\theta \pm f(\theta, \Delta^2)/(\sqrt{2} \sin 2\theta) \quad (19a)$$

$$Z_{\pm} = \sqrt{1 + [g_{\pm}(\theta, \Delta^2)]^2} \quad (19b)$$

Substituting the previous equations into Eq. (7) we obtain (remember that for qubit systems  $k = \pm 1/2$ ):

$$H(\theta, \Delta^2) = e^{-2\Delta^2} \sin^2 2\theta, \quad (20)$$

which reaches the maximum for  $\theta = \pi/4$ : the best states for phase estimation also in the presence of phase noise are the equatorial ones, i.e., the states laying in  $x$ - $y$  plane of the Bloch sphere.

Since the Bures metrics, and then, the Bures distance [19, 20, 21, 22, 23, 24, 25] between states, is proportional to the QFI [17, 26], this result can be easily understood under a geometrical point of view. If we choose two states, one equatorial and the other not, and shift them by the same amount  $\phi$ , then the *distance* on the Bloch sphere between the initial states and the shifted counterparts is larger for the equatorial states (see Fig. 3) which, in turn, allow better estimation.

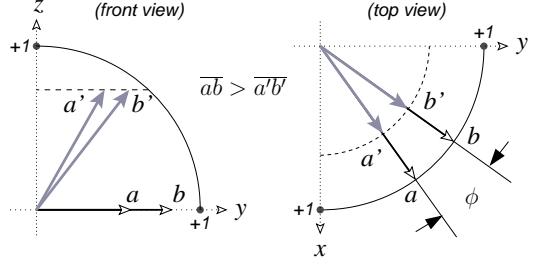


FIG. 3: The distance  $\overline{ab}$  on the Bloch sphere between two equatorial states separated by a phase shift  $\phi$  is larger than the distance  $\overline{a'b'}$  between non equatorial states separated by the same amount  $\phi$ . The same conclusion holds in the presence of phase noise: best estimation is achieved involving equatorial states.

The optimal quantum estimator leading to (20) can be written as [17]:

$$O_{\phi} = \phi \mathbb{I} + \frac{L_{\phi}}{H(\theta, \Delta^2)}, \quad (21)$$

where we introduced the symmetric logarithmic derivative:

$$\partial_{\phi} \varrho_{\phi} = \frac{L_{\phi} \varrho_{\phi} + \varrho_{\phi} L_{\phi}}{2}, \quad (22)$$

and  $\varrho_{\phi}$  is the state of the qubit initially prepared in a pure state with Bloch vector  $\mathbf{r}$  given by (13), shifted by the application of the unitary transformation  $U_{\phi} = \exp(-i\phi\sigma_+\sigma_-) = \exp[-\frac{i}{2}\phi(\sigma_z + \mathbb{I})]$  and degraded by the evolution through the noisy environment:

$$\varrho_{\phi} = \begin{pmatrix} \cos^2 \theta & e^{-i\phi - \Delta^2} \cos \theta \sin \theta \\ e^{i\phi - \Delta^2} \cos \theta \sin \theta & \sin^2 \theta \end{pmatrix}. \quad (23)$$

One finds:

$$L_{\phi} = i \frac{2g_+g_-(g_+ - g_-)(\lambda_- - \lambda_+)}{Z_+^2 Z_-^2 (\lambda_+ + \lambda_-)} (\sigma_+ e^{i\phi} - \sigma_- e^{-i\phi}), \quad (24)$$

where  $\lambda_{\pm}$ ,  $g_{\pm} \equiv g_{\pm}(\theta, \Delta^2)$  and  $Z_{\pm}$  are given in (16) and (19), respectively. If we choose  $\theta = \pi/4$ , Eq. (24) reduces to:

$$L_{\phi} = i e^{-\Delta^2} (\sigma_+ e^{i\phi} - \sigma_- e^{-i\phi}). \quad (25)$$

### III. PHASE-SHIFT ESTIMATION BY SPIN MEASUREMENT

Let us now consider a realistic scenario where, in order to estimate  $\phi$ , we measure the spin in a generic direction in the plane, i.e. the observable

$$\Theta_{\alpha} = \sigma_x \cos \alpha + \sigma_y \sin \alpha, \quad (26)$$

whose eigenvectors  $|\Sigma_{\pm}(\alpha)\rangle$ ,  $\Theta_{\alpha}|\Sigma_{\pm}(\alpha)\rangle = \pm|\Sigma_{\pm}(\alpha)\rangle$ , are:

$$|\Sigma_{\pm}(\alpha)\rangle = \frac{1}{\sqrt{2}} (e^{-i\alpha}| + 1/2\rangle \pm | - 1/2\rangle). \quad (27)$$

For the qubit of Eq. (23), the probabilities to obtain the outcomes  $\pm 1$  given the phase shift  $\phi$  read:

$$P_{\Delta^2}(\pm 1|\phi) = \text{Tr}[|\Sigma_{\pm}(\alpha)\rangle\langle\Sigma_{\pm}(\alpha)|\varrho_{\phi}] \quad (28)$$

$$= \frac{1}{2} \left[ 1 \pm e^{-\Delta^2} \cos(\alpha - \phi) \sin 2\theta \right], \quad (29)$$

and the expectation value is:

$$\langle\Theta_{\alpha}\rangle = \text{Tr}[\Theta_{\alpha}\varrho_{\phi}] = e^{-\Delta^2} \cos(\alpha - \phi) \sin 2\theta. \quad (30)$$

The corresponding Fisher information turns out to be:

$$F(\phi, \Delta^2) = \sum_{k=\pm 1} P_{\Delta^2}(k|\phi) [\partial_{\phi} \ln P_{\Delta^2}(k|\phi)]^2 \quad (31)$$

$$= \frac{e^{-2\Delta^2} \sin^2(\alpha - \phi) \sin^2 2\theta}{1 - e^{-2\Delta^2} \cos^2(\alpha - \phi) \sin^2 2\theta}, \quad (32)$$

which is plotted in Fig. 4 in the case of equatorial probe states ( $\theta = \pi/4$ ) as a function of  $\delta = \alpha - \phi$  and different values of  $\Delta^2$ .

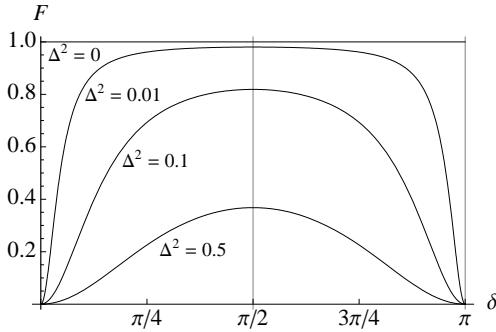


FIG. 4: Plot of the Fisher information  $F(\phi, \Delta^2)$  in the case of equatorial qubit probe states ( $\theta = \pi/4$ ) as a function of  $\delta = \alpha - \phi$  and different values of  $\Delta^2$ .

As we can see, in the absence of noise ( $\Delta^2 = 0$ ) one has  $F = 1, \forall \alpha, \phi$ , i.e., the Fisher information is equal to the QFI  $H$  in Eq. (20) [27, 28]. When noise affects the propagation, the maximum of  $F$ , that corresponds to the QFI  $H$ , is achieved for  $\delta = \alpha - \phi = \pi/2$ , while goes to zero as  $\alpha - \phi = k\pi, k \in \mathbb{N}$ . These results can be better understood considering the sensitivity of the measurement (actually this is the square of the sensitivity):

$$\mathcal{S} = \frac{\text{Var}[\Theta_{\alpha}]}{(\partial_{\phi}\langle\Theta_{\alpha}\rangle)^2} = \frac{1 - \langle\Theta_{\alpha}\rangle^2}{(\partial_{\phi}\langle\Theta_{\alpha}\rangle)^2}, \quad (33)$$

that is the ratio between the fluctuations of  $\langle\Theta_{\alpha}\rangle$  and how  $\langle\Theta_{\alpha}\rangle$  varies with respect to  $\phi$ ;  $\sqrt{\mathcal{S}}$  represents the smallest change of  $\phi$  that can be detected with our measurement (up to the statistical scaling, of course). In the present case:

$$\mathcal{S}(\phi, \Delta^2) = \frac{1 - e^{-2\Delta^2} \cos^2(\alpha - \phi) \sin^2 2\theta}{e^{-2\Delta^2} \sin^2(\alpha - \phi) \sin^2 2\theta}, \quad (34)$$

which is just the inverse of Eq. (32): the maximum of  $F$  (maximum information) corresponds to the case of maximum sensitivity (minimum of  $\mathcal{S}$ ). If  $\Delta^2 = 0$  one finds that  $\text{Var}[\Theta_{\alpha}]$

and  $(\partial_{\phi}\langle\Theta_{\alpha}\rangle)^2$  are always equal, no matter the values of  $\alpha$  and  $\phi$ . When noise is acting, the maximum of  $F$  at  $\delta = \pi/2$  corresponds to the minimum of  $\mathcal{S}(\phi, \Delta^2)$ , this fact can be also understood by geometrical means addressing the special case of  $\alpha = 0$  ( $\Theta_0 = \sigma_x$ ). In this case the result of the measurement carried out on the probe is just the projection onto the  $x$ -axis: for a fixed change  $d\phi$ , the change of  $\langle\Theta_0\rangle$  at  $\phi = 0$  is smaller than the one at  $\phi = \pi/2$  (see Fig. 5).

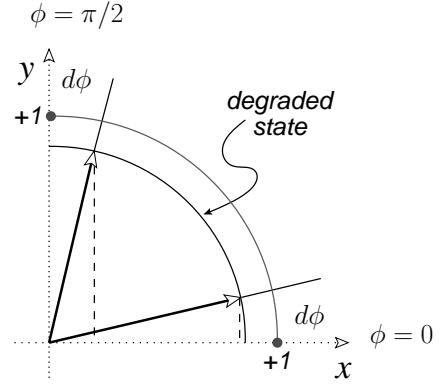


FIG. 5: Sensitivity of  $\sigma_x$  measurement to a small change  $d\phi$  in the phase-shift. The projections onto  $x$ -axis are the expectations  $\langle\Theta_0\rangle$ , with  $\Theta_0 = \sigma_x$ . For a fixed change  $d\phi$ , the change of  $\langle\Theta_0\rangle$  at  $\phi = 0$  is smaller than the one at  $\phi = \pi/2$ . For the sake of simplicity we sketch only the upper right quarter of the Bloch sphere  $x$ - $y$  equatorial plane.

The Fisher information depends on the actual, unknown value of  $\phi$ . However, we can perform an *adaptive*, two-step method to achieve the QFI. During the first step, we use a small amount of data to obtain a rough estimate  $\tilde{\phi}$  of the phase shift, then, at the second step, we *tune*  $\Theta_{\alpha}$  according to the transformation  $\alpha \rightarrow \tilde{\phi} + \pi/2$ : the (eventual) repetition of these two steps allows to reach the QFI limit. The same result can be obtained by fixing the measurement at a chosen  $\alpha$  and tuning the probe state by applying a suitable rotation.

#### IV. PHASE-SHIFT ESTIMATION FOR POLARIZATION ENCODED OPTICAL QUBITS

In our proposal the qubit state corresponds to the polarization degree of freedom of a coherent state. We refer to  $|H\rangle$  and  $|V\rangle$  respectively as horizontal and vertical polarization. Initially, we set the polarization at  $|+\rangle = \frac{1}{\sqrt{2}}(|H\rangle + |V\rangle)$  (equatorial state) and apply the phase shift with eventual Gaussian noise (as we have shown in section II this is equivalent to a non dissipative phase noise). Then we measure  $\sigma_x$ , which corresponds to inserting a half wave plate (HWP) at  $22.5^\circ$  in front of a polarizing beam splitter (PBS) and to record the number of counts  $n_+$  and  $n_-$  at the two outputs, which correspond to the outcomes  $+1$  and  $-1$ , respectively. Once a data sample  $X$  has been acquired, one needs to give an estimation of the phase shift. Here we consider two strategies: one based on the inversion of the probability functions, whose uncertainties come from the error propagation; the other applies the Bayes

theorem to retrieve a prior distribution of phase shift, knowing this distribution we can calculate all the desired moments.

### A. Estimation by inversion

If we send  $M$  copies of our state, then we can write:

$$P_{\Delta^2}(\pm 1|\phi) = \frac{n_{\pm}}{n_+ + n_-}, \quad (35)$$

and, by inverting Eq. (29), after some algebra we obtain the following expression for the phase shift estimator:

$$\phi_{\text{inv}} = \arccos \left( \frac{n_+ - n_-}{n_+ + n_-} e^{\Delta^2} \right). \quad (36)$$

The uncertainty in the estimation is thus given by:

$$\text{Var}[\phi_{\text{inv}}] = \left( \frac{\partial \phi_{\text{inv}}}{\partial n_+} \right)^2 \sigma^2(n_+) + \left( \frac{\partial \phi_{\text{inv}}}{\partial n_-} \right)^2 \sigma^2(n_-), \quad (37)$$

where  $\sigma^2(n_{\pm})$  are the fluctuations of the numbers of outcomes  $n_{\pm}$ . It is worth noting that at each shot, *i.e.*, for each sent copy of the coherent state, the detected number of photons fluctuates according to the laws of quantum mechanics.

### B. Bayesian estimation

Given the data sample  $X = \{x_1, x_2, \dots, x_N\}$ , where  $x_k \in \{-1, +1\}$ ,  $\forall k$ , we can define the *sample* probability:

$$P(X|\phi) = \prod_{k=1}^N P_{\Delta^2}(x_k|\phi), \quad (38)$$

that is the probability to obtain the whole data sample  $X$  given the unknown phase  $\phi$ . By means of the Bayes theorem one can write the *a posteriori* probability [31]:

$$P(\phi|X) = \frac{1}{\mathcal{N}} \prod_{k=1}^N P_{\Delta^2}(x_k|\phi), \quad \mathcal{N} = \int_{\Phi} d\phi P(\phi|X), \quad (39)$$

where  $\Phi$  is the parameter space. The probability (39) is the probability distribution of  $\phi$  *given* the data sample  $X$ . The Bayesian estimator are thus obtained as:

$$\phi_B = \int_{\Phi} d\phi \phi P(\phi|X), \quad (40)$$

$$\text{Var}[\phi_B] = \int_{\Phi} d\phi (\phi - \phi_B)^2 P(\phi|X). \quad (41)$$

Bayesian estimators are known to be asymptotically optimal, namely, they allow one to achieve the Cramér-Rao bound as the size of the data sample increases [17, 29, 30, 31, 32].

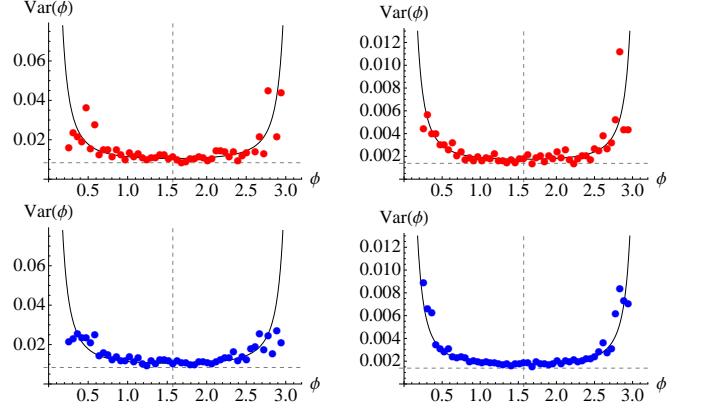


FIG. 6: Variances from simulated Monte Carlo experiments performed with  $M = 60$  copies of the input state and  $\Delta = 0.34$ . The plots on the top refer to the inversion method, the bottom ones to Bayesian analysis. For the plots on the left we chose a coherent state with an average number of photons equal to  $\bar{n} = 2$ , for those on the right we set  $\bar{n} = 12$ . The solid line is the Cramér-Rao bound  $1/(FM\bar{n})$ . The dashed horizontal line is the quantum Cramér-Rao bound  $1/(HM\bar{n})$ . Notice the asymptotic optimality of the Bayesian estimation.

### C. Monte Carlo simulated experiments

Fig. 6 shows the variances of the estimated phases from Monte Carlo simulated experiments as function of the unknown phase shift. The simulations have been performed sending  $M = 60$  copies of the input coherent state and choosing  $\Delta = 0.34$ . The plots on the top refer to the inversion method, the bottom ones to Bayesian analysis. We investigated two different cases: low energy (average number of photons equal to  $\bar{n} = 2$ , left plots), and high energy ( $\bar{n} = 12$ , rights plots). The solid line is the Cramér-Rao bound obtained from Eq. (34), while the dashed horizontal line is the quantum Cramér-Rao bound both rescaled by the effective number of measurements  $N = M\bar{n}$ . We can see that in the low energy regime there is a more evident deviation from the expected behavior (solid line) at the extremes of the phase interval  $(0, \pi)$ . This is the manifestation of a systematic error due to phase window sampling, *i.e.* the tails of the Gaussian distributions, simulating the phase-noise, is truncated at the boundary and refolded inside the interval. Notice the asymptotic optimality of the Bayesian estimation when the energy, and, thus, the number of events increases. It is also useful to observe how the fluctuations of the variances obtained by using the Bayesian method are less than the other ones, which actually depend on the fluctuations of the average number of photons. A further investigation of Fig. 6 leads us to conclude that both the methods are robust over a large interval of phases: for the considered examples the variance is almost constant over the phase interval  $(0.5, 2.5)$ .

## V. EXPERIMENTAL SETUP AND RESULTS

The experimental demonstration of our scheme is based on a KDP crystal, that allows both the manipulation of the optical qubit polarization and the simulation of a phase diffusion environment. In Fig. 7 we sketch the experimental setup.

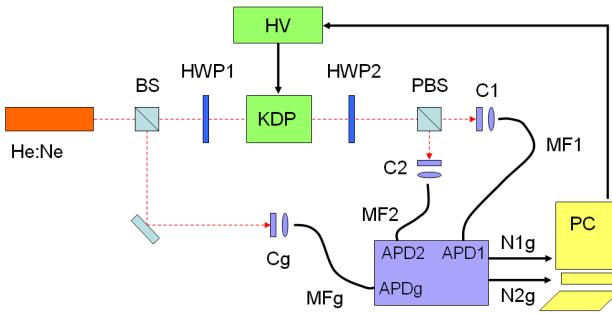


FIG. 7: Sketch of the experimental setup to estimate the phase of a polarization qubit in the presence of non dissipative noise. The qubit state is set to  $|+\rangle$  by inserting a half-wave plate (HWP1) along the path of the He:Ne laser output, while a KDP crystal, driven by a high voltage generator (HV), is used both to add the phase shift  $\phi$  and to simulate the noise. A second half-wave plate @  $22.5^\circ$  (HWP2) together with a polarizing beam splitter (PBS) implement the measurement (the couplers, C, and the multi-mode fibers, MF, are used to bring the signals to the avalanche photodiodes, APD). Finally, the contemporary counts  $N_{1g}$  and  $N_{2g}$  coming from the APD1 and APD2 and from the APDg, i.e., the gate channel signal coming from the beam splitter (BS) placed in front of the laser source, are acquired by a PC module.

A linearly polarized He:Ne laser (Thorlabs HRP120) generates coherent states that are prepared in the initial qubit state  $|+\rangle = \frac{1}{\sqrt{2}}(|H\rangle + |V\rangle)$  by means of a first half-waveplate (HWP1). Then the qubits pass through a KDP crystal driven by a stabilized high voltage generator (HV), up to a maximum of 6 kV. The application of an electrostatic field  $E_z$  along the  $z$  axis of the KDP (that is also the direction of propagation of the qubit), a phase shift  $\phi(E_z)$  is introduced between the qubit polarization components along  $x$  and  $y$ . The crystal is oriented with the  $x$  axis parallel to the horizontal polarization and tilted around the vertical axis  $y$  in order to avoid the multiple internal reflections superposed with the main beam, that are removed by a pin hole.

The detection system consists of a second half-waveplate (HWP2), a polarizing beam splitter (PBS), absorption filters and three detectors. The HWP2 plate is set at  $22.5^\circ$  with respect to the horizontal axis and the PBS selects the polarization. Light signals are focused into multi-mode fibers (MF) and sent to the detectors APD1 and APD2. The detectors are single photon counting modules (SPCM) based on avalanche photodiodes operated in Geiger mode with passive quenching. For the coincidence counts an electronic circuit based on AND gates is used. A 50:50 Beam Splitter (BS) is placed before the HWP1 plate to add a gate channel (g) for the counting measurement. The aim is to make acquisition with a low number

of counts maintaining an high signal to noise ratio  $N_i/N_{i,dc}$  ( $i = 1, 2$ ). This is the reason for the coincidence counting, according to the formula ( $i = 1, 2$ ):

$$N_{ig} = (N_i + N_{i,dc})N_g\Delta t = N_iN_g\Delta t + N_{i,dc}N_g\Delta t = N_{ig,true} + N_{ig,dc}, \quad (42)$$

where  $N_{ig}$ ,  $N_i$  and  $N_g$  are respectively the coincidence and direct counts on the detectors APD*i* and APDg,  $N_{ig,dc}$  and  $N_{i,dc}$  are respectively the coincidence and direct dark counts, and  $\Delta t$  is the coincidence time window of the electronic counting module ( $\Delta t = 90$  ns).  $N_{ig}$  could be as low as we want by selecting the time window of the acquisition and the ratio  $N_{ig,true}/N_{ig,dc} = N_i/N_{i,dc}$  remain constant. We measure about  $10^5$  counts/s on the detector APDg and a maximum of about  $9 \times 10^4$  counts/s on the other two but we can adjust these direct counts changing filters or the detector voltages above breakdown. The dark counts are below 200 counts/s. Each acquisition is taken with a 10 ms time window in order to have  $\bar{n} \simeq 10$  in the contemporary counts. Note that the signal to noise ratio remain  $\simeq 10^3$ . The rising and falling time of the high voltage generator are measured around 200 ms so the waiting time between two subsequent measurement is set at 240 ms.

We made  $M=60$  acquisitions for each phase-shift  $\phi$ . Notice that each acquisition corresponds to a different random phase distributed according to a Gaussian distribution centered on  $\phi$  and with standard deviation  $\sigma^2(\phi) = \Delta^2/2$  [see Eq. (3)]. In our implementation for each acquisition we send a different voltage to the KDP crystal according to a proper calibration curve  $\phi(V)$ .

The results are shown in Fig. 8, where we plot the theoretical Cramér-Rao bound (solid line) and the experimental variances obtained using the inversion method (red squares) and the Bayesian analysis (blue circles). The Cramér-Rao bound is quickly reached by Bayesian estimation, even for a relatively small number of measurements ( $N = M\bar{n}$  in the figure, with  $M = 60$  and  $\bar{n} \simeq 10$ ).

### A. Adaptive method for optimality

As we have seen Bayesian estimation allows us to reach the Cramér-Rao bound for any orientation of the measurement angle. On the other hand, as we pointed out in the section III, optimal estimation, i.e. with precision at quantum Cramér-Rao limit is achieved only if  $\phi = \pi/2$ . In this Section we show that ultimate precision can be always achieved upon the application of an adaptive method [32, 33, 34]. In the present case, we proceed as follows: i) we start our estimation obtaining a first value  $\phi_1$  of the unknown phase; ii) we set  $\alpha$  of the measured observable (26) to the value  $\alpha_1 = \pi/2 + \phi_1$  (this can be also obtained by adding the phase shift  $\alpha'_1 = \pi/2 - \phi_1$  to the qubit itself) so the new phase is near to  $\pi/2$ . One can also repeat many times this procedure in order to refine the estimation. However, as one can see in Figs. 6 and 8, the optimal condition is almost achieved after the first step and the variance is nearly constant around the optimality region. We

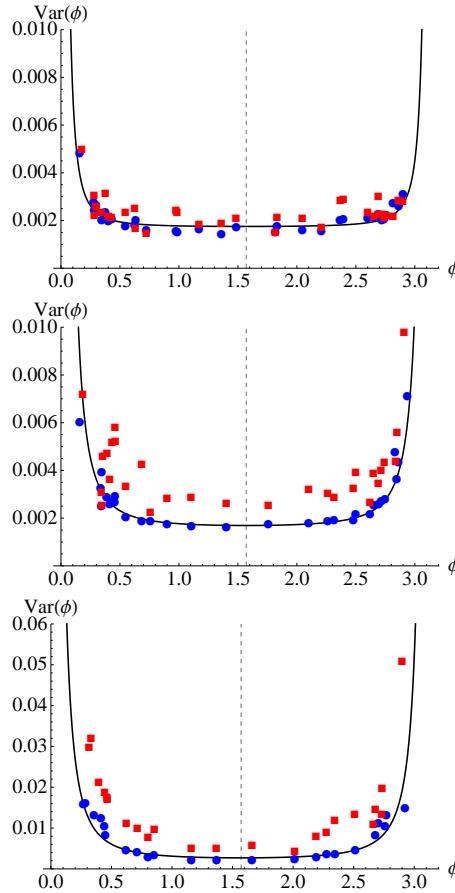


FIG. 8: (Color online) Variances from experiments performed with  $M = 60$  copies of the input state and, from top to bottom,  $\Delta = 0.13 \pm 0.02$ ,  $\Delta = 0.24 \pm 0.03$  and  $\Delta = 0.48 \pm 0.05$ . (Red) Squares refer to inversion method, (blue) circles to Bayesian analysis. We also have, from top to bottom,  $\bar{n} = 9.83 \pm 0.93$ ,  $\bar{n} = 11.06 \pm 0.43$  and  $\bar{n} = 9.78 \pm 0.95$ . The solid line is the Cramér-Rao bound  $1/(FM\bar{n})$ .

report the results in Table I, where one can read the step number of the adaptive method, the corresponding variance and the estimated phase. Note that already at the second step optimal estimation has been achieved, that is the experimental variances are close to the optimal variance given by the quantum Cramér-Rao bound.

## VI. CONCLUSIONS

We have demonstrated an optimal estimation scheme for the phase-shift imposed to a polarization encoded optical

qubit in the presence of phase diffusion. Our scheme is based on polarization measurement assisted by Bayesian estimation and allows to achieve the ultimate quantum bound to precision using a limited number of measurements. In turn, Bayes estimator is known to be asymptotically unbiased, but for practical implementation is of interest to evaluate quantitatively how many measurements are needed to achieve the asymptotic region. Our results indicate that Bayesian inference represents a useful tool for phase estimation. An adaptive method to achieve optimal estimation in any working regime, i.e. for any value of the unknown phase-shift, has been also analyzed in details and experimentally demonstrated. As a future perspective we foresee the possibility to investigate the effects of different kind of noises and to employ entangled probes to increases the overall stability of the estimation procedure, a results which has been theoretically established for the ideal (i.e without noise) case [31].

step	$\text{Var}(\phi)$	$\phi^{(\text{est})}$
1	$4.21 \times 10^{-3}$	$0.33 \pm 0.06$
2	$2.49 \times 10^{-3}$	$0.22 \pm 0.05$
3	$2.52 \times 10^{-3}$	$0.19 \pm 0.05$
4	$2.59 \times 10^{-3}$	$0.20 \pm 0.05$

TABLE I: Phase estimation using the adaptive method (see the text for details). The actual value of the phase is  $\phi = 0.17 \pm 0.01$ . The amplitude of the phase noise is  $\Delta = 0.46 \pm 0.06$  and the energy of the coherent state is  $\bar{n} = 10.97 \pm 0.67$ . For each step we used  $M = 55$  repetitions for the estimation. For these values of the parameters the quantum Cramér-Rao bound, evaluated from Eq. (20) and rescaled by the number of measurements  $N = M\bar{n}$ , is equal to  $\text{Var}(\phi) \simeq 2.52 \times 10^{-3}$

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